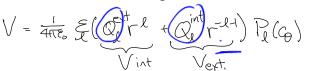
Section 6.3 - Auxiliary Field H

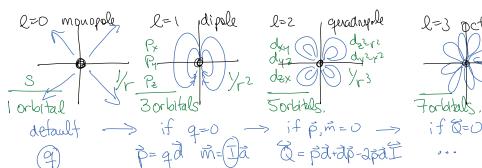
* reminder - Multipoles: gerneral solution to $\nabla^2 V = 0$ with azimuthal symmetry



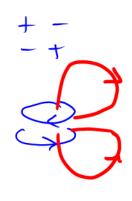
$$\frac{1}{4\pi\xi_0} \rightarrow \frac{\nu_0}{4\pi}$$

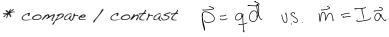
$$V \rightarrow \bigcup$$

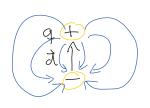
* magnetic multipoles

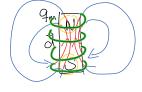


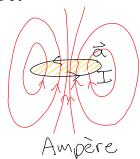
l=3 octupole

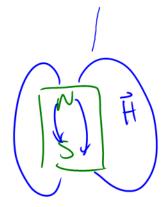












DuFay

Gilbert

* polarization $\tilde{\mathcal{P}}$, magnetization $\tilde{\mathbb{M}}$ = dipole density

- ~ to get effective charge / current distribution:
 - a) expand V into dipole potential
 - b) integrate V due to dipole density field

 - c) $\frac{1}{2} \frac{1}{2} = -\frac{1}{2} \frac{1}{2} \frac{1}{2$

A=器感染 P>dp=Pdt m>dm=mh 0 = - 8.P J= 8xM

~ can also expand magnetic scalar potential to get "magnetic pole density" directly analogous with electric charge

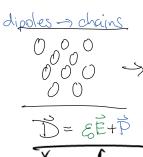
Pm=-5.M

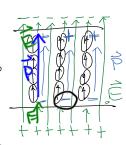
~ given permanent \vec{P} or \vec{M} , use ρ_0 ρ_m or \vec{J}_m to calculate fields

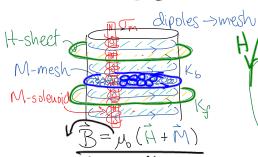


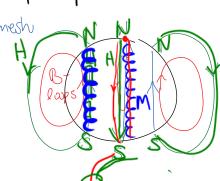
м。(Ĥ+M)

* polarization chain; magnetization solenoid and mesh.







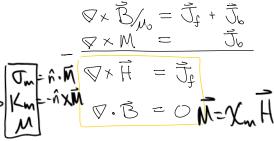


* source fields D, H only include free charge/current as sources

$$\nabla \cdot \mathcal{C} \stackrel{?}{=} P + P$$

$$+ \frac{\nabla \cdot \vec{P}}{\nabla \cdot \vec{P}} = -P$$

$$\nabla \cdot \vec{P} = P$$



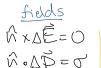
$$\vec{B} = \mu_0 (\vec{H} + \vec{M}) = \mu \vec{H}$$

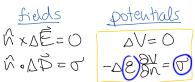
$$\mu = \mu_0 \mu_1 = \mu_0 (l + \chi_m)$$

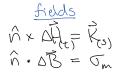
* boundary conditions: \$\top \hat{n} \Delta \tag{\tau} \data

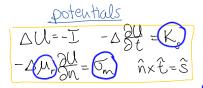
~ don't double count bound charge: Kb, Tm, Ur

all account for the same thing!









Om=- n.M

- * three ways to solve similar magnetic boundary value problems:

 - a) use Gilbert "pole density" pm of explicitly b) use Ampere "bound current" Jo explicitly
 - c) absorb magnetization into "permeability" U
- ~ See Example #1
- ~ See Example #2
- ~ See Example #3





* Example 1: Magnetic pole density of

$$\nabla \times \vec{H} = \vec{J}_f = \vec{0} \qquad \Rightarrow H = -\nabla U$$

$$\nabla \cdot \vec{B} = \nabla \cdot \mu_o (\vec{H} + \vec{M}) = 0 \Rightarrow -\nabla^2 U = \rho_m \text{ no } \mu_o!$$
where
$$\rho_m = -\nabla \cdot \vec{M} \qquad \nabla_m = \hat{n} \cdot \vec{M} = M_o \cos \theta$$

$$U_{1} = \sum_{k=0}^{\infty} \left[A_{k} (7a)^{k} + B_{k} (\%)^{k+1} \right] P_{k} (c_{0})$$

$$U_{2} = \sum_{k=0}^{\infty} \left[C_{k} (7a)^{k} + D_{k} (\%)^{k+1} \right] P_{k} (c_{0})$$

$$U_1 = \frac{1}{3} M_0 r \cos \theta = \frac{1}{3} M_0 z$$

$$\overrightarrow{H}_1 = -\overrightarrow{M}_3$$

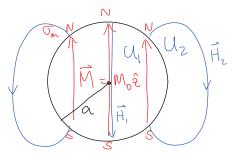
$$\overrightarrow{B}_1 = \mu_0 (\overrightarrow{H} + \overrightarrow{M}) = \frac{3}{3} \mu_0 \overrightarrow{M}$$

$$\lim_{\alpha \to 0} : \overrightarrow{E} = \underbrace{\frac{\partial \vec{p}}{\partial t}}_{\text{See prob 3.42 p 157}} - \underbrace{\frac{1}{3}\vec{p}}_{\text{See prob 3.42 p 157}} - \underbrace{\frac{1}{3}\vec{m}}_{\text{See prob 3.42 p 157}} - \underbrace{\frac{1}$$

$$\vec{H} = \frac{Q\vec{m}}{4\pi r^3} - \frac{1}{3}\vec{m} S^3(\vec{r})$$

See prob 3.42 p 15+
$$\vec{B} = \frac{\mu_0 \vec{Q} \vec{n}}{4\pi \vec{r}^3} + \frac{2}{3} \vec{p} \vec{S}^3(\vec{r})$$
See prob 3.42 p 15+
$$\vec{B} = \frac{\mu_0 \vec{Q} \vec{n}}{4\pi \vec{r}^3} + \frac{2}{3} \mu_0 \vec{m} \vec{S}^3(\vec{r})$$
See prob 5.59 p 253

note closed B= No(A+M) lines of flux.



* Example 2: Bound current density \vec{J}_{ij}

 $\nabla \times \vec{H} = \vec{J}_{tot} = 0$ except on $\partial \rightarrow \vec{H} = -\nabla U$ discontinuity of U from $\nabla \cdot \vec{D} = -\nabla \cdot \mu_0 \nabla U = 0$ $\rightarrow \nabla^2 U = 0$ Kb on the boundary $\vec{J}_b = \nabla \times \vec{M} = 0$ $\vec{K}_b = -\hat{n} \times \vec{M} = M_0 \sin \theta \hat{\phi}$ (treat K_b as a free current)

BC's: A gu = 0 - 1 gu = Ks

 $U_{1} = \underbrace{20}_{0} \left[A_{e} \left(\frac{7}{4} \right)^{l} + B_{e} \left(\frac{9}{4} \right)^{l+1} \right] P_{e} (c_{0})$ $U_{2} = \underbrace{20}_{0} \left[C_{e} \left(\frac{7}{4} \right)^{l} + D_{e} \left(\frac{9}{4} \right)^{l+1} \right] P_{e} (c_{0})$

 $BC^{\pm}: -\frac{\partial U_2}{\partial r} + \frac{\partial U_1}{\partial r} = \underbrace{\tilde{\mathcal{E}}}_{\text{end}} \left[-D_{\ell} \left(l + l \right) \left(\frac{-l}{\alpha} \right) + A_{\ell} \left(l \right) \left(\frac{-l}{\alpha} \right) \right] P_{\ell} \left(c_{\Theta} \right) = O \quad D_{\ell} = \frac{-l}{l+l} A_{\ell}$

 $BC^{*}2: -\frac{\partial U_{2}}{\partial \theta}\Big|_{a} + \frac{\partial U_{1}}{\partial \theta}\Big|_{a} = \underbrace{\mathcal{E}}_{e=0} \Big[-D_{e} + A_{e} \Big] + \underbrace{\mathcal{E}}_{a} \Big[(C_{0}) (-\sin \theta) = M_{0} \sin \theta$

note: $P_{\omega}(c_0)$ form a basis, and $P_{\omega}(x) = 1$ like the RHS. so $(-D_1 + A_1) \cdot \frac{-1}{\alpha} = (\frac{1}{2} - 1) A_{1/\alpha} = M_0$ $A_1 = -\frac{9}{3} a M_0 = -2D_1$ Km=n×M

 $U_1 = -\frac{1}{3} M_0 = U_2 = \frac{1}{3} M_0 \frac{\cos \theta}{r^2} = \frac{\overrightarrow{m} \cdot \overrightarrow{r}}{4\pi r^3}$ $\overrightarrow{B}_1 = y_0 \overrightarrow{H}_1 = y_0 \frac{2}{3} \overrightarrow{M}$ $\overrightarrow{B}_2 = \frac{Q \overrightarrow{m}}{4\pi r^3}$

* Note: $\vec{B} = \mu_0 \vec{H}$, not $\mu_0(\vec{H} + \vec{M})$ inside because we replaced \vec{M} with its effective current distribution K_b * thus \vec{H} is different from $\vec{E}_x^{\pm}1$, but \vec{B} is still the same.

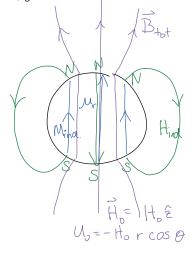
* Example 3: Magnetic permeability (linear homogeneous isotropic material)

Permeable Sphere in an external constant field | +

M=XnA encapsulated in B= MA see Ex4.7 p186

$$\nabla x \vec{H} = \vec{J}_{tot} = 0$$
 everywhere $\rightarrow \vec{H} = -\nabla U$
 $\nabla \cdot \vec{B} = -\nabla \cdot \mu \nabla U = 0 \rightarrow \nabla^2 U = 0$
 $BC_S': U(\omega) \rightarrow -H_0 Z, \Delta U = 0, \Delta \mu \partial U = 0$

$$U_{1} = \mathcal{E}_{0} \left(A_{\ell} \left(\frac{r_{0}}{a} \right)^{\ell} + B_{\ell} \left(\frac{\alpha}{r} \right)^{\ell+1} \right) P_{\ell} \left(c_{0} \right)$$



BC#0: $\lim_{r\to\infty} U_2(r) = \mathop{\mathcal{E}}_{s,s} C_e(7a)^l P_e(c_0) = -H_0 r C_0$ $C_1 = -H_0 a$ we exclude all multipoles except l=1, since the source \widehat{H}_s is pure dipole.

$$BC^{\pm}$$
: $U_{2}|_{\alpha} - U_{1}|_{\alpha} = (-H_{0}\alpha + D_{1} - A_{1}) = 0 \Rightarrow D_{1} = A_{1} + H_{0}\alpha$

$$BC^{*}2: -\frac{\partial U_{2}}{\partial r} + \mu \frac{\partial U_{1}}{\partial r} = (H_{0} + D_{1} \cdot 2/a + \mu_{r} A_{1/a}) \cos \theta = 0$$

$$3H_0 + (2+\mu_r)A_1/a = 0$$
 $A_1/a = \frac{-3}{2+\mu_r}H_0$ $D_1/a = \frac{-1+\mu_r}{2+\mu_r}H_0 = \frac{x_m}{2+\mu_r}H_0$

$$U_1 = U_0 + \frac{\chi_m}{a+\mu_r} H_0 Z = \frac{-3}{a+\mu_r} H_0 Z \qquad \hat{H}_1 = \frac{3}{a+\mu_r} \hat{H}_0$$

$$U_2 = U_0 + \frac{\chi_m}{a+\mu_r} H_0 \alpha^3 \frac{\cos \Theta}{r^2} \quad \text{where } U_0 z - H_0 Z$$

* compare with Ex#1: U,= U0+ 3MZ U2= U0+ 303M cos 0

$$\vec{N} = \frac{3 \, \text{Km}}{2 + \text{Mr}} \vec{H}_0 = \text{Km} \left(\frac{3}{2 + \text{Mr}} \right) \vec{H}_0 = \text{Km} \vec{H}_1$$
 as dictated by $\vec{M} = \text{Km} \vec{H}_1$

$$\vec{B}_{1} = \frac{3\mu}{2\pi} \vec{H}_{0} \qquad \vec{B}_{2} = \mu_{0} \vec{H}_{2} \qquad \text{where } \vec{m} = \frac{4\pi\alpha^{3} \cdot \frac{3m}{2\pi} H_{0}}{2\pi} H_{0} = \mu_{0} (\vec{H}_{0} + \frac{\vec{D}\vec{m}}{2\pi r^{3}}) \qquad \vec{Q}\vec{m} = 3\hat{r}\hat{r} \cdot \vec{m} - \vec{m}$$