Symposium in honor of Keh-Fei Liu on the occasion of his 60th Birthday



- A good physicist
 - wide knowledge, deep intuition, full of innovative ideas, up-todate in theory and experiment
- Visionary
 - For example: implementation of the overlap operator (reached a record 180 MeV pion mass on a 3.2 fm lattice, exposed ghosts in the chiral regime)
- Quality over quantity
 - "A study of pentaquarks on the lattice with overlap fermions", Kentucky group, PRD70 (2004).
 - no evidence of bound state, consistent with KN scattering states
 - volume dependence of the spectral weight (1/V). $G(t) = w \exp(-m t)$
 - described as the "most credible lattice calculation" at the plenary Koff Joint Corr, UKY, 1

Background Field Calculations in Lattice QCD

Frank Lee, GWU

polarQCD Collaboration <u>http://eagle.phys.gwu.edu/~fxlee/polarQCD.html</u>

- Physics motivation
- Background field method
- Some results
 - magnetic moments
 - magnetic polarizability
 - electric polarizability
- Outlook





If the proton and neutrons in this picture were 10 cm across, then the quarks and electrons would be less than 0.1 mm in size and the entire atom would be about 10 km across.

A journey deep into the heart of the atom

Quarks and Strong Interactions

	name	mass (GeV/c)	EM charge (e)
u	up	0.003	+2/3
d	down	0.006	-1/3
S	strange	0.08	-1/3
с	charm	1.3	+2/3
b	bottom	4.5	-1/3
t	top	174	+2/3



But, it is a long, hard struggle from quarks and gluons to ...

The Particle Zoo (excitation spectrum of QCD)

Baryons (3 quarks)

Mesons (quark-antiquark)

0	P.,	****	A(1232)	Paa	****	Λ	P.,,	****	5+	P.,	****	=0	P.,	****		LIGHT UN	FLAVO RED		STRAM	NGE	BOTT (8 -	FOM +11
p n	P.,	****	A(1600)	P	***	A(1405)	S	****	<u>5</u> 0	P.,	** **		P.,	** **		1º(J ^{ec})		$I^{G}(J^{PC})$	(<i>v</i> = 10.0 0	(J ^P)	(# <u>-</u>	1 ^e (f ^e)
 N(1440)	P.,	****	A(1600)	r 33 C	****	A(1520)	Д.,	****	5-	P.,	** **	=(1530)	P	** **	• =	1-(0-)	• x2(1670)	$1^{-}(2^{-+})$	• K±	1/2(0-)	• B±	1/2(0-)
N(1520)	D	****	$\Delta(1020)$ $\Lambda(1700)$	531	****	A(1600)	D 03	***	Σ(1385)	P	** **	=(1600)	, 13	*	• π ⁰	1-(0-+)	• d(1680)	0-(1)	• K ⁰	1/2(0-)	• B ⁰	1/2(0-)
N(1520)	C 13	****	∆(1700)	D ₃₃	+	A(1(70)	F 01	++++	Σ(1490)	/ 13	*	=(1600)		***	• n • \$(600)	0+(0++) 0+(0++)	 ρ₂(1690) 	$1^+(3^{})$ $1^+(1^{})$	• K ⁰	1/2(0-)	 B±/B⁰ ADMI ADMI 	AT URE
N(1535)	511	++++	$\Delta(1750)$	P ₃₁	*	V(10/0)	S ₀₁	****	Z(1400)		**	=(1090)		***	• A(770)	1+(1)	ap(1700)	$1^{-}(2^{++})$	Kt(800)	1/2(0+)	MIXTURE	oaryon Aco-
N(1650)	511	****	$\Delta(1900)$	S ₃₁	**	A(1690)	D_{03}	****	2 (1500)	_	**	=(1820)	D_{13}	***	 ω(782) 	0-(1)	 f₀(1710) 	0+(0++)	 K*(892) 	1/2(1-)	V _{c0} and V _{u0} C Elements	K M Matrix
N(1675)	D ₁₅	****	$\Delta(1905)$	F_{35}	****	A(1800)	S ₀₁	***	Σ(1580)	D_{13}	**	=(1950)		***	 η'(958) η'(958) 	$0^+(0^-+)$ $0^+(0^++)$	η(1760)	0+(0-+)	 K₁(1270) 	1/2(1+)	• B*	$1/2(1^{-})$
N(1680)	F ₁₅	****	$\Delta(1910)$	P ₃₁	****	A(1810)	P ₀₁	***	Σ(1620)	S ₁₁	**	Ξ(2030)		***	• a (980)	$1^{-}(0^{+})$	• 1(1800) 6(1810)	$0^{+}(2^{++})$	 K₁(1400) K[*](1410) 	1/2(1+)	B'j(5732)	3(5,)
N(1700)	D ₁₃	***	$\Delta(1920)$	P ₃₃	***	A(1820)	F_{05}	****	Σ(1660)	P ₁₁	***	Ξ(2120)		*	• d(1020)	0-(1)	 \$\phi_1(1850)\$ 	0-(3)	 K[*]₀(1430) 	1/2(0+)	BOTTOM,	STRANGE
N(1710)	P ₁₁	***	$\Delta(1930)$	D_{25}	***	A(1830)	D_{05}	****	Σ(1670)	D_{13}	** **	Ξ(2250)		**	 b₁ (1170) 	0-(1+-)	$\eta_2(1870)$	0+(2-+)	 K[*]₂(1430) 	1/2(2+)	(B = ±1,	5 = #1]
N(1720)	P13	****	$\Delta(1940)$	D22	*	A(1890)	Pm	****	Σ(1690)		**	Ξ(2370)		**	 b₁ (1235) a₁ (1260) 	$1^{+}(1^{+})$ $1^{-}(1^{+})$	ρ(1900) 6(1910)	1+(1) 0+(2++)	K(1460)	1/2(0-)	• B ^o ₂	0(0-)
N(1900)	P.,	**	A(1950)	- 33 Faa	****	A(2000)	. 03	*	Σ(1750)	S11	***	=(2500)		*	 6(1270) 	$0^{+}(2^{+}^{+})$	 fs(1950) 	$0^{+}(2^{++})$	R ₂ (1580) K(1630)	1/2(27)	B*,(5850)	n(??)
N(1990)	E.,	**	A(2000)	F 37	**	A(2020)	E	*	$\Sigma(1770)$	P11	*	-(f₁(1285) 	0+(1++)	$\rho_{1}(1990)$	1+(3)	K ₁ (1650)	1/2(1+)	BOTTOM	LIADMED
N(2000)	F.,	**	$\Delta(2000)$ $\Lambda(2150)$	r35 c	*	A(2100)	6	****	$\Sigma(1775)$	D.,,	** **	0-		** **	 η(1295) η(1300) 	$0^+(0^-+)$ $1^-(0^-+)$	 f₂(2010) f₂(2020) 	$0^+(2^{++})$ $0^+(0^{++})$	 K*(1680) 	1/2(1-)	(B = C	= ±1}
N(2000)	/15 D	**	A(2000)	531	÷.	A(2110)	0 ₀₇	***	T(1040)	D 15	*	0(2250)-		***	• a ₂ (1320)	1-(2++)	• a (2040)	$1^{-}(4^{++})$	 K₂(1770) K²(1780) 	1/2(2-)	• B [±] _c	0(0-)
N(2080)	D ₁₃	+	$\Delta(2200)$	637	Ť.	71(2110)	r ₀₅		Z(1040)	r ₁₃	**	0(2200)-		**	 6(1370) 	0+(0++)	 \$4(2050) 	0+(4++)	 K₂(1820) 	$1/2(2^{-})$	ci	Ē
N(2090)	511	Ţ	$\Delta(2300)$	H39	**	A(2325)	D ₀₃	*	Z (1000)	P ₁₁	****	0(2470)-		**	h ₁ (1380)	?=(1+=)	π ₂ (2100)	$1^{-}(2^{-+})$	K(1830)	1/2(0-)	 η_c(15) 	0+(0-+)
N(2100)	P_{11}	*	$\Delta(2350)$	D_{35}	*	A(2350)	H ₀₉	***	2 (1915)	P15	****	32(2410)			• n(1405)	0+(0-+)	f(2150)	$0^{+}(0^{+})^{+}(2^{+})^{+}$	K [*] ₀ (1950)	1/2(0+)	 J/\$\phi\$(15) 	0-(1)
N(2190)	G_{17}	****	$\Delta(2390)$	F ₃₇	*	A(2585)		**	Σ(1940)	D_{13}	***	a+		****	 f₁(1420) 	0+(1++)	p(2150)	1+(1)	• Kt(2045)	1/2(2+)	• $\chi_{c0}(1P)$	$0^+(0^++)$ $0^+(1^++)$
N(2200)	D ₁₅	**	$\Delta(2400)$	G39	**				Σ(2000)	S ₁₁	*	/1' c		****	 ω(1420) 	0-(1)	\$(2200)	$0^{+}(0^{++})$	K ₂ (2250)	1/2(2-)	b _c (1P)	2 ² (2 ⁷⁷)
N(2220)	H ₁₉	****	$\Delta(2420)$	H _{3.11}	****				Σ(2030)	F ₁₇	****	$\Lambda_{c}(2593)^{+}$		***	5(1430) • a(1450)	$1^{-}(0^{+}^{+})$	13(2220)	or 4 + +)	K ₃ (2320)	1/2(3+)	 χ_{c2}(1P) 	0+(2++)
N(2250)	G19	****	$\Delta(2750)$	413	**				Σ(2070)	F ₁₅	*	$\Lambda_{c}(2625)^{+}$		***	• A(1450)	1+(1)	η(2225)	0+(0-+)	K [*] ₂ (2380)	$1/2(5^{-})$ $1/2(4^{-})$	nd(25)	0+(0 - +)
N(2600)	h 11	***	A(2050)	Kerr	**				Σ(2080)	P ₁₃	**	$A_{c}(2765)^{+}$		*	 η(1475) 	0+(0-+)	$\rho_{2}(2250)$	1+(3)	K(3100)	21072	• \$(25) • \$(3770)	0 (1)
N(2700)	K1 12	**	±(2,500)	r3,15					Σ(2100)	G17	*	$\Lambda_{c}(2880)^{+}$		**	• f ₀ (1500) 6(1510)	$0^{+}(0^{+}^{+})$	 5(2300) 6(2300) 	$0^+(2^{++})$ $0^+(4^{++})$	CHAR	MED	¢(3836)	0-(2)
(,			Q(1540)+		***				Σ(2250)		***	Σ-(2455)		** **	 f^r₂(1525) 	$0^{+}(2^{+}^{+})$	 f₂(2340) 	$0^{+}(2^{++})$	(C=	±1)	X(3872)	??(??)
			A(1960)		*				Σ(2455)		**	Σ.(2520)		***	6(1565)	0+(2++)	$\rho_{5}(2350)$	1+(5)	• D±	1/2(0-)	 	0-(1)
			Ψ(1000)		+				5(2620)		**	=+		***	h ₁ (1595)	$0^{-(1+-)}$	a₂(2450)	$1^{-}(6^{++})$	• D ⁰	1/2(0-)	 ¢(4415) 	0-(1)
									T(2000)		*	=0		***	a,(1640)	$1^{-}(1^{+})$	10(2510)	0.(0)	 D*(2007)^o D*(2010)[±] 	1/2(1-)	5	5
									Z (3000)		÷	-c -++		+++	f2(1640)	0+(2++)	OTHER L	IGHT	 D₁ (2420)^o 	1/2(1+)	======================================	0+ /0 - +)
									2(3170)		*	-c		***	• 12 (1645)	0+(2-+)	Further States		$D_1(2420) \pm$	1/2(??)	• T(15)	0-(1)
												=_c		***	• u(1650) • ua(1670)	0-(1)			 D⁺₂(2460)⁰ D⁺₂(2460)[±] 	$1/2(2^+)$	 χ_M(1P) 	0+(0++)
												$\Xi_{c}(2645)$		***		* (*)			 D₂(240)[±] D[*](2640)[±] 	1/2(2?)	• $\chi_{\Delta 1}(1P)$ • $\chi_{\omega}(1P)$	$0^+(1^++)$ $0^+(2^++)$
												$\Xi_{c}(2790)$		***					- ()	-/ -(. /	• T(25)	0-(1)
												$\Xi_{c}(2815)$		***					CHARMED, 1 (C=5=	STRANGE = ±1]	 χ_{Δ0}(2P) 	0+(0 + +)
												Ω^0		***					• D±	0(0-)	 χ_{b1}(2P) (2P) 	0+(1++)
												c							• D**	0(??)	 <i>χ_M</i>(2P) <i>τ</i>(35) 	0-(1)
												=+		*					 D[*]_{sJ}(2317)[±] 	0(0+)	 T(45) 	0-(1)
												-cc							 D_{sJ}(2460)* D_s(2536)* 	0(1+)	 T(10860) 	0-(1)
												<u>/0</u>		***					 D₁₂(2573)[±] 	0(??)	• 7(11020)	0 (1)
												=0 =-		*							NO N-q q CA	NDIDATES
												-b, -b									NON-99 CAN	DIDATES

 $L_{QCD} = \frac{1}{2} \operatorname{Tr} F_{\mu\nu} F^{\mu\nu} + \overline{q} (\gamma^{\mu} D_{\mu} + m_{q}) q$

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Expansion of the Universe

After the Big Bang, the universe expanded and cooled. At about 10⁻⁶ second, the universe consisted of a soup of quarks, gluons, electrons, and neutrinos. When the temperature of the Universe, T_{universe}, cooled to about 10¹² K, this soup coalesced into protons, neutrons, and electrons. As time progressed, some of the protons and neutrons formed deuterium, helium, and lithium nuclei. Still later, electrons combined with protons and these low-mass nuclei to form neutral atoms. Due to gravity, clouds of atoms contracted into stars, where hydrogen and helium fused into more massive chemical elements. Exploding stars (supernovae) form the most massive elements and disperse them into space. Our earth was formed from supernova debris.



2006 Nobel Prize in Physics

• Study of the very small is closely related to the study of the very big.

Quantum Chromodynamics (QCD)

--- the fundamental theory of the strong interaction (in terms of quarks and gluons)

$$S_{QCD} = \int d^4x \left[\frac{1}{2} \operatorname{Tr} F_{\mu\nu} F^{\mu\nu} + \overline{q} (\gamma^{\mu} D_{\mu} + m_q) q \right]$$

All physics is computed from path integrals

$$\left\langle \hat{O} \right\rangle \equiv \frac{\int DG_{\mu} Dq D\bar{q} e^{-S_{QCD}} \hat{O}(G_{\mu}, q, \bar{q})}{\int DG_{\mu} Dq D\bar{q} e^{-S_{QCD}}}$$



which can be evaluated on a space-time lattice.

Path Integral Method (Richard Feynmann)



• Applicable to any problem that can be cast into the form

$$\langle O \rangle \equiv \frac{\int Dx \ e^{-S} O(x)}{\int Dx \ e^{-S}}$$

with an action, for example:

$$S = \int_{t_1}^{t_2} dt \left\{ \frac{1}{2} m \dot{x}^2(t) + V[x(t)] \right\}$$

- Successfully used in
 - statistical physics
 - quantum chemistry
 - condensed matter physics
 - biological physics
 - quantum field theories (QED,QCD, ...) $_{X_1}$
 - and more



Background-Field Calculations



Polarizabilities

Interaction energy of a hadron in the presence of external electromagnetic fields:

$$H = -\vec{\mu} \cdot \vec{B} - \frac{1}{2} \alpha E^2 - \frac{1}{2} \beta B^2$$
$$-\frac{1}{2} \gamma_{E1} \sigma \cdot \vec{E} \times \dot{\vec{E}} - \frac{1}{2} \gamma_{M1} \sigma \cdot \vec{B} \times \vec{E}$$
$$+ \gamma_{E2} \sigma_i E_{ij} B_j - \gamma_{M2} \sigma_i B_{ij} E_j$$
$$-\frac{1}{12} \alpha_{E2} E_{ij}^2 - \frac{1}{12} \beta_{M2} B_{ij}^2 + \cdots$$

Time and spatial derivatives : $\dot{E} = \frac{\partial E}{\partial t}$, $E_{ij} = \frac{1}{2} (\nabla_i E_j + \nabla_j E_i)$, etc

Probe of internal structure of the system in increasingly finer detail. μ, α, β :

static bulk response

others :

spatial and time resolution

Compton Scattering

Low-energy expansion of real Compton scattering amplitude on the nucleon



$$\begin{aligned} T_{\gamma N} &= A_1(\omega,\theta) \,\vec{\epsilon}' \cdot \vec{\epsilon} + A_2(\omega,\theta) \,\vec{\epsilon}' \cdot \hat{k} \,\vec{\epsilon} \cdot \hat{k}' + i \,A_3(\omega,\theta) \,\vec{\sigma} \cdot (\vec{\epsilon}' \times \vec{\epsilon}) + i \,A_4(\omega,\theta) \,\vec{\sigma} \cdot (\hat{k}' \times \hat{k}) \,\vec{\epsilon}' \cdot \vec{\epsilon} \\ &+ i \,A_5(\omega,\theta) \,\vec{\sigma} \cdot \left[(\vec{\epsilon}' \times \hat{k}) \,\vec{\epsilon} \cdot \hat{k}' - (\vec{\epsilon} \times \hat{k}') \,\vec{\epsilon}' \cdot \hat{k} \right] + i \,A_6(\omega,\theta) \,\vec{\sigma} \cdot \left[(\vec{\epsilon}' \times \hat{k}') \,\vec{\epsilon} \cdot \hat{k}' - (\vec{\epsilon} \times \hat{k}) \,\vec{\epsilon}' \cdot \hat{k} \right] \end{aligned}$$

$$\begin{split} A_{1}(\omega,\theta) &= -Z^{2} \frac{e^{2}}{M_{N}} + \frac{e^{2}}{4M_{N}^{3}} \left(\mu^{2}(1+\cos\theta) - Z^{2}\right) (1-\cos\theta) \,\omega^{2} + 4\pi(\alpha+\beta\cos\theta)\omega^{2} + \mathcal{O}(\omega^{4}) \,, \\ A_{2}(\omega,\theta) &= \frac{e^{2}}{4M_{N}^{3}} (\mu^{2} - Z^{2})\omega^{2}\cos\theta - 4\pi\beta\omega^{2} + \mathcal{O}(\omega^{4}) \,, \\ A_{3}(\omega,\theta) &= \frac{e^{2}\omega}{2M_{N}^{2}} \left(Z(2\mu-Z) - \mu^{2}\cos\theta\right) + 4\pi\omega^{3}(\gamma_{1} - (\gamma_{2}+2\gamma_{4})\cos\theta) + \mathcal{O}(\omega^{5}) \,, \\ A_{4}(\omega,\theta) &= -\frac{e^{2}\omega}{2M_{N}^{2}}\mu^{2} + 4\pi\omega^{3}\gamma_{2} + \mathcal{O}(\omega^{5}) \,, \\ A_{5}(\omega,\theta) &= \frac{e^{2}\omega}{2M_{N}^{2}}\mu^{2} + 4\pi\omega^{3}\gamma_{4} + \mathcal{O}(\omega^{5}) \,, \\ A_{6}(\omega,\theta) &= -\frac{e^{2}\omega}{2M_{N}^{2}}Z\mu + 4\pi\omega^{3}\gamma_{3} + \mathcal{O}(\omega^{5}) \,, \end{split}$$

polarizabilities: α , β , γ_1 , γ_2 , γ_3 , γ_4

Experimental Information on Polarizabilities

- Proton electric polarizability (α_p) is around 12 in units of 10⁻⁴ fm³.
- Proton magnetic polarizability (β_p) is around 2 in units of 10⁻⁴ fm³.
- α_n is about the same as α_p
- β_n is about the same as β_p
- Experiments are under way or planned for other polarizabilities at electron accelerators around the world

Polarizabilities from ChPT

	CHF	РТ	Δ -pole		Dispe	rsion R	elation	s	
	(leading order)			A_2^{as}		excitat	tions +	$\operatorname{ons} + A_1^{\operatorname{as}}$	
					pro	ton	ne	utron	
	π^{0}	loop			HDT	SAID	HDT	SAID	
α_E		12.3				11.9 ^{<i>a,c</i>}		13.3 ^{b,c}	
β_M		1.2	12.0			$1.9^{a,c}$		$1.8^{b,c}$	
$(10^{-4}{\rm fm^3})$									
$\alpha_{E\nu}$		2.2				-3.8		-2.4	
$\beta_{M\nu}$		3.5	5.3			9.1		9.2	
α_{E2}		20.7	0.2			27.5		27.2	
β_{M2}		-8.9				-22.4		-23.5	
$(10^{-4}{\rm fm}^5)$									
γ_{E1}	$11.3\tau_{3}$	-5.5		$11.2\tau_{3}$	-4.5	-3.4	-5.5	-5.6	
γ_{M1}	$-11.3\tau_{3}$	-1.1	4.0	$-11.2\tau_{3}$	3.4	2.7	3.4	3.8	
γ_{E2}	$-11.3\tau_{3}$	1.1	-0.75	$-11.2\tau_{3}$	2.3	1.9	2.6	2.9	
γ_{M2}	$11.3\tau_{3}$	1.1		$11.2\tau_{3}$	-0.6	0.3	-0.6	-0.7	
γ		4.4	-4.8		-0.6	-1.5	-0.2	-0.4	
γ_{π}	$-45.3\tau_{3}$	4.4	4.8	$-45.0\tau_{3}$	10.8	7.8^{d}	12.1	13.0	
$(10^{-4}{\rm fm}^4)$									

Babusci et al, PRC58, 1013 (1998)

Polarizabilities on the Lattice

Measure mass shifts in progressively-small external electric and magnetic fields, specially designed to isolate them:

$$H = -\vec{\mu} \cdot \vec{B} - \frac{1}{2} \alpha E^2 - \frac{1}{2} \beta B^2$$
$$-\frac{1}{2} \gamma_{E1} \sigma \cdot \vec{E} \times \dot{\vec{E}} - \frac{1}{2} \gamma_{M1} \sigma \cdot \vec{B} \times \dot{\vec{B}}$$
$$+ \gamma_{E2} \sigma_i E_{ij} B_j - \gamma_{M2} \sigma_i B_{ij} E_j$$
$$-\frac{1}{12} \alpha_{E2} E_{ij}^2 - \frac{1}{12} \beta_{M2} E_{ij}^2 + \cdots$$

Small field expansion:

$$\delta m(B) = m(B) - m(0) = c_1 B + c_2 B^2 + c_3 B^3 + c_4 B^4 + \cdots$$

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Introduction of an external electromagnetic field on the lattice

• Minimal coupling in the QCD covariant derivative in Euclidean space $D_{\mu} \rightarrow \partial_{\mu} + gG_{\mu} + qA_{\mu}$

Recall that SU(3) gauge field is introduced by the link variables

$$U_{\mu}(x) = \exp(iagG_{\mu})$$

• It suggests multiplying a U(1) phase factor to the links

$$U'_{\mu}(x) = \exp(iaqA_{\mu})U_{\mu}$$

• This should be done in two places where the Dirac operator appears: both in the dynamical gauge generation and quark propagator generation

For Example

• To apply magnetic field **B** in the z-direction, one can choose the 4-vector potential

$$A_{\mu} \equiv (\phi, \vec{A}) = (0, 0, Bx, 0)$$

$$B = \nabla \times A$$
$$\vec{E} = -\nabla \phi - \frac{\partial \vec{A}}{\partial t}$$

then the y-link is multiplied by a x-dependent phase factor $U_v \rightarrow \exp(iqaBx)U_v$



• To apply electric field **E** in the x-direction, one can choose the 4-vector potential $A_{\mu} = (0, Et, 0, 0)$

then the x-link is multiplied by a t-dependent phase factor

$$U_x \rightarrow \exp(iqaEt)U_x$$

Background Field Method

Wilson action, 6³x10 lattice L=1 fm, pion mass about 1400 and 900 MeV, 9 configurations. done on a VAX

- "Lattice quantum-chromodynamics calculation of some baryon magnetic moments", Bernard, Draper, Olynyk, PRL49 (1982) 1076; NPB220 (1983) 508
- "Dependence of lattice hadron masses on external magnetic fields", Rubinstein, Solomon and Wittlich, NPB457 (1995) 577
- "A study of hadron electric polarizability in quenched lattice QCD", Fiebig, Wilcox, Woloshyn, NPB324, 47 (1989)
- "Electric Polarizability of Neutral Hadrons from Lattice QCD", Christensen, Wilcox, Lee, Zhou, PRD72, 034503 (2005)
- "Baryon magnetic moments in the background field method", Lee, Kelly, Zhou, Wilcox, PLB627, 71 (2005)
- "Magnetic polarizability of hadrons from lattice QCD in the background field method", Lee, Zhou, Wilcox, Christensen, PRD73, 034503 (2006)
- "Electricmagnetic and spin polarizabilitites in lattice QCD", Detmold, Tiburzi, Walker-Loud, PRD73 (2006) 114505
- "Neutron electric dipole moment with external electric field method in lattice QCD", Shintani et al, CP-PACS collaboration, PRD75, 034507 (2007)

Computational Demands

• Consider quark propagator generation

$$\frac{\int DG_{\mu} \det(\mathbf{D} + m_q) e^{-S_c} (\mathbf{D} + m_q)^{-1}}{\int DG_{\mu} \det(\mathbf{D} + m_q) e^{-S_c}}$$

$$D_{\mu} \rightarrow \partial_{\mu} + gG_{\mu} + qA_{\mu}$$

 $U_v \rightarrow \exp(iqaBx)U_v$

- For each value of external field, a new dynamical ensemble is needed that couple to u-quark (q=1/3) and d- and s-quark (q=-2/3).
 - quark propagator is then computed on the ensembles with matching values
- Cost is much reduced if no field is applied in the vacuum: any gauge ensemble can be used to compute valence quark propagators.

Lattice details

- Standard Wilson gauge action
 - -24^4 lattice, β =6.0 (or a ≈ 0.1 fm)
 - 150 configurations
- Standard Wilson fermion action
 - $-\kappa$ =0.1515, 0.1525, 0.1535, 0.1540, 0.1545, 0.1555
 - Pion mass about 1015, 908, 794, 732, 667, 522 MeV
 - Strange quark mass corresponds to κ =0.1535 (or m_{π}~794 MeV)
 - Fermion boundary conditions: periodic in y and z, fixed in x and t
 - Source location (t,x,y,z)=(2,12,1,1)
- The following 5 dimensionless numbers $\eta \equiv qBa^2 = +0.00036$, -0.00072, +0.00144, -0.00288, +0.00576 correspond to 4 small B fields

 $eBa^2 = -0.00108, 0.00216, -0.00432, 0.00864$ for both u and d (or s) quarks.

- Small in the sense that the mass shift is only a fraction of the proton mass: $\mu B/m ~ 1 to 5% at the smallest pion mass. In physical units, B ~ 10^{13} Tesla.$



 $U_{v} \rightarrow \exp(iqaBx)U_{v}$

What about boundary conditions?

• On a finite lattice with periodic boundary conditions, to get a constant magnetic field, B has to be quantized by

$$qBa^2 = \frac{2\pi n}{N_x}, n = 1, 2, 3, \cdots$$

to ensure that the magnetic flux through plaquettes in the x-y plane is constant.



 $U_{v}^{B} = \exp(iqa^{2}Bx)$

- But, for $N_x=24$ and 1/a=2 GeV, the lowest allowed field would give the proton a mass shift of about 500 MeV, which is unacceptably large (proton is severely distorted). So we have to abandon the quantization condition, and work with much smaller fields.
- To minimize the boundary effects, we work with fixed b.c. in xdirection, so that quarks originating in the middle of the lattice has little chance of propagating to the edge.

A computational trick

- We generate two sets of quark propagators, one with the original set of fields, the other with the fields reversed.
- The mass shift in the presence of small fields is

 $\delta m(B) = m(B) - m(0) = c_1 B + c_2 B^2 + c_3 B^3 + c_4 B^4 + \cdots$

- At the cost of a factor of two,
 - by taking the average, $[\delta m(B) + \delta m(-B)]/2$, we get the leading quadratic response with the odd-powered terms eliminated. (magnetic polarizability)
 - by taking the difference, [δm(B) δ m(-B)]/2, we get the leading linear response with the even-powered terms eliminated.
 (magnetic moment)
- Our calculation is equivalent to 11 mass spectrum calculations.
 - 5 original fields, 5 reversed, plus the zero-field to set the baseline

Magnetic Moments



Magnetic moment

- For a Dirac particle of spin s in small fields, $E_{\pm} = m \pm \mu B$ where upper sign means spin-up and lower sign spindown, and $\mu = g \frac{e}{2m} s$
- g factor is extracted from

$$g = m \frac{(E_+ - m) - (E_- - m)}{eBs}$$

• Look for the slope (g-factor) on the straight line of the form $\delta m = g(eB)$

Proton mass shifts

$\delta m = g(eB)$



• We use the 2 smallest fields to fit the line.



Neutron mass shifts



Proton and neutron magnetic moments



• To one meson loop, χ PT predicts $\mu = \mu_0 + c_1 m_{\pi} + c_2 m_{\pi}^2 \log m_{\pi}^2 + c_3 m_{\pi}^2 + \cdots$

but only applicable in small mass region.

• Encapsulating form (Leinweber, Lu and Thomas, PRD60 (1999) 034014) μ_0

$$\mu = \frac{\mu_0}{1 + \alpha m_\pi + \beta m_\pi^2}$$

Pade ansatz

• For small mass,

$$\mu = \mu_0 [1 - \alpha m_\pi + (\alpha^2 - \beta) m_\pi^2 + \cdots]$$

• For large mass,

$$\mu = \frac{\mu_0}{\beta m_\pi^2} \left(1 - \frac{\alpha}{\beta m_\pi} + \cdots \right) \propto \frac{1}{m_q}$$

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Octet Sigma magnetic moments



Delta magnetic moments



Proton and Δ^+ magnetic moments



Curvatures expected from ChPT.

Magnetic moments

Table 1

The computed magnetic moments for the baryon octet and decuplet in nuclear magnetons (μ_N) as a function of the pion mass. The extrapolated values are based on Eq (10). The experimental values are taken from the PDG [16]

K m (CiaM)	0.1515	0.1525	0.1535 0.704	0.1540 0.722	0.1545 0.667	0.1555	Extrap.	Expt.
m_{π} (GeV)	1.015	0.908	V.794	V.732	0.007	0.522		
p	1.48(3)	1.63(4)	1.80(6)	1.90(8)	2.02(10)	2.34(17)	3.04(6)	2.79
п	-0.94(2)	-1.02(2)	-1.13(3)	-1.18(4)	-1.25(5)	-1.39(8)	-1.84(3)	-1.91
Σ^+	1.57(4)	1.70(5)	1.85(6)	1.93(7)	2.02(8)	2.23(11)	2.87(3)	2.45
Σ^0	0.52(1)	0.55(1)	0.59(2)	0.61(2)	0.63(2)	0.67(3)	0.76(1)	0.65
Σ^{-}	-0.54(4)	-0.60(4)	-0.68(6)	-0.73(6)	-0.78(7)	-0.92(9)	-1.48(5)	-1.16
S^0	-1.07(4)	-1.10(4)	-1.13(4)	-1.15(4)	-1.17(5)	-1.21(7)	-1.37(1)	-1.25
Σ^{-}	-0.76(6)	-0.77(7)	-0.77(8)	-0.77(8)	-0.77(9)	-0.78(11)	-0.82(1)	-0.65
Λ	-0.59(2)	-0.60(2)	-0.60(2)	-0.61(2)	-0.61(2)	-0.62(3)	-0.70(1)	-0.61
Δ^{++}	3.35(7)	3.68(9)	4.06(12)	4.28(14)	4.50(16)	4.92(28)	5.24(18)	4.52(1.00)
Δ^+	1.55(4)	1.67(6)	1.77(8)	1.80(10)	1.80(11)	1.64(24)	0.97(8)	
Δ^0	-0.002(0)	-0.003(0)	-0.004(1)	-0.005(1)	-0.007(1)	-0.011(6)	-0.035(2)	
Δ^{-}	-1.58(4)	-1.73(5)	-1.89(7)	-1.98(8)	-2.07(10)	-2.34(17)	-2.98(19)	
$\Sigma *^+$	1.40(5)	1.56(6)	1.72(8)	1.80(10)	1.86(11)	1.88(17)	1.27(6)	
$\Sigma *^0$	-0.13(1)	-0.09(1)	-0.03(0)	-0.01(0)	0.03(0)	0.08(2)	0.33(5)	
$\Sigma *^-$	-1.68(5)	-1.76(7)	-1.83(9)	-1.87(10)	-1.89(11)	-1.93(16)	-1.88(4)	
$\Xi *^0$	-0.09(2)	-0.06(1)	-0.022(5)	-0.002(1)	0.018(6)	0.05(3)	0.16(4)	
5*-	-0.59(7)	-0.61(8)	-0.62(9)	-0.62(10)	-0.63(10)	-0.63(11)	-0.62(1)	

F.X. Lee, R. Kelly, L. Zhou, W. Wilcox, Phys. Lett. B 627, 71 (2005)

Our new results on vector meson g factors (preliminary)





$$< \rho | J_{\mu} | \rho > \sim \{G_{C}(q^{2}), G_{M}(q^{2}), G_{Q}(q^{2})\}$$

hep-lat/0703014, Adelaide group

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Polarizabilities





Neutron Mass Shift in Electric Field



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Electric Polarizability of neutron



Electric Polarizabilities of Neutral Particles

TABLE II: The electric polarizabilities from the calculation with Wilson action using six κ values.

The units of the electric polarizability are 10^{-4} fm³. The pion masses were fit on time steps 11 - 13

from the propagator origin and are given in GeV.

Christensen, Wilcox, Lee, Zhou, Phys.Rev. D72 (2005) 034503

ĸ	0.1515	0.1525	0.1535	0.1540	0.1545	0.1555	
m_{π}	$1.000\pm.005$	$0.895 \pm .006$	$0.782\pm.006$	$0.721\pm.006$	$0.657 \pm .007$	$0.512\pm.007$	
Mesor	1						fit range
ρ^0	5.0 ± 0.5	4.8 ± 0.7	4.4 ± 0.9	4.0 ± 1.2	3.5 ± 1.6	2.8 ± 4.8	13 - 15
$K^{\ast 0}$	1.2 ± 0.2	1.2 ± 0.2	1.3 ± 0.3	1.3 ± 0.3	1.3 ± 0.4	1.2 ± 0.6	10-12
Baryo	n octet						
n	7.9 ± 0.5	8.6 ± 0.7	9.5 ± 1.0	10.2 ± 1.4	10.8 ± 1.9	10.6 ± 5.7	14-16
Σ^0	7.7 ± 0.7	8.1 ± 0.9	8.8 ± 1.3	9.3 ± 1.5	10.0 ± 1.9	11.5 ± 4.0	14-16
Λ^0_o	8.2 ± 0.7	8.8 ± 0.8	9.7 ± 1.1	10.2 ± 1.2	10.8 ± 1.6	13.2 ± 3.2	14-16
Ξ^0	9.3 ± 0.9	9.7 ± 1.0	10.0 ± 1.2	10.2 ± 1.4	10.3 ± 1.6	10.1 ± 2.3	14-16
Baryo	n decuplet						
Δ^0	1.7 ± 0.1	1.6 ± 0.2	1.5 ± 0.3	1.5 ± 0.4	1.6 ± 0.6	2.3 ± 1.2	9-11
$\Sigma^{\ast 0}$	1.7 ± 0.2	1.6 ± 0.2	1.5 ± 0.4	1.5 ± 0.4	$1.6~\pm~0.5$	2.0 ± 0.9	9-11
Ξ^{*0}	1.7 ± 0.2	1.6 ± 0.3	1.5 ± 0.4	1.5 ± 0.4	1.5 ± 0.5	1.6 ± 0.8	9-11

Neutron Mass Shifts in Magnetic Field



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 βB^2

 $\Delta m =$

Magnetic Polarizability of the Nucleon



Magnetic Polarizabilities: baryon octet

TABLE III. The calculated magnetic polarizabilities for the octet baryons as a function of the pion mass from the Wilson action. The pion mass is in GeV and the magnetic polarizability is in 10^{-4} fm³. The time window from which each polarizability is extracted is given in the last column. The errors are statistical.

κ	0.1515	0.1525	0.1535	0.1540	0.1545	0.1555	fit range
m_{π}	1.000	0.895	0.782	0.721	0.657	0.512	
Р	0.09 ± 0.29	0.14 ± 0.37	0.26 ± 0.48	0.40 ± 0.56	0.64 ± 0.67	2.36 ± 1.20	12-14
n	9.4 ± 0.2	10.4 ± 0.3	11.6 ± 0.4	12.3 ± 0.5	13.4 ± 0.6	17.0 ± 1.1	12 - 14
Σ^+	$\textbf{-0.15}\pm0.36$	0.09 ± 0.42	0.24 ± 0.50	0.40 ± 0.56	0.61 ± 0.64	1.60 ± 1.00	12-14
Σ^0	8.0 ± 0.3	8.5 ± 0.3	9.1 ± 0.4	9.6 ± 0.5	10.1 ± 0.6	11.9 ± 0.9	12-14
Σ^{-}	$\textbf{-11.8}\pm0.3$	-12.6 \pm 0.3	-13.6 \pm 0.4	$\textbf{-14.2}\pm0.4$	$\textbf{-14.7}\pm0.4$	-16.1 \pm 0.5	12-14
Ξ^0	10.7 ± 0.3	11.3 ± 0.4	11.9 ± 0.4	12.3 ± 0.4	12.8 ± 0.5	13.9 ± 0.7	12-14
Ξ^-	-12.6 ± 0.3	-13.1 \pm 0.3	-13.8 \pm 0.4	$\textbf{-14.1}\pm0.4$	-14.6 \pm 0.4	-15.6 \pm 0.5	12 - 14
Λ^8	9.1 ± 0.4	9.9 ± 0.5	10.8 ± 0.6	11.4 ± 0.7	12.1 ± 0.8	14.0 ± 1.2	12-14
Λ^C	9.3 ± 0.4	10.2 ± 0.5	11.2 ± 0.6	11.9 ± 0.7	12.6 ± 0.9	15.0 ± 1.4	12-14
Λ^S	3.6 ± 0.1	$\textbf{3.4}\pm \textbf{0.2}$	$\textbf{3.3}\pm\textbf{0.2}$	$\textbf{3.3}\pm\textbf{0.3}$	3.2 ± 0.4	3.2 ± 1.0	5-7

F.X. Lee, L. Zhou, W. Wilcox, J. Christensen, Phys. Rev. D73 (2006) 034503

Magnetic Polarizabilities: baryon decuplet

TABLE IV. The calculated magnetic polarizabilities for the decuplet baryons as a function of the pion mass from the Wilson action. The pion mass is in GeV and the magnetic polarizability is in 10⁻⁴ fm³. The time window from which each polarizability is extracted is given in the last column. The errors are statistical.

κ	0.1515	0.1525	0.1535	0.1540	0.1545	0.1555	fit range
m_{π}	1.000	0.895	0.782	0.721	0.657	0.512	
Δ^{++}	-39.9 ± 0.7	-43.9 ± 0.8	-48.7 ± 1.0	-51.5 ± 1.1	-54.7 ± 1.3	-63.1 ± 1.9	10-12
Δ^+	-2.5 ± 0.2	-3.0 ± 0.3	-3.5 ± 0.5	-3.9 ± 0.6	-4.3 ± 0.7	-5.1 ± 1.1	10 - 12
Δ^0	7.6 ± 0.2	8.2 ± 0.3	8.8 ± 0.4	9.2 ± 0.5	9.6 ± 0.6	$10.9 \pm\ 1.0$	10 - 12
Δ^-	-10.1 ± 0.2	-11.1 \pm 0.2	-12.4 ± 0.2	-13.1 \pm 0.3	-14.0 ± 0.3	-16.2 ± 0.5	10 - 12
Σ^{*+}	-2.9 ± 0.3	-3.2 ± 0.4	-3.6 ± 0.5	-3.9 ± 0.6	-4.2 ± 0.6	-5.1 ± 0.9	10-12
Σ^{*0}	7.9 ± 0.2	8.4 ± 0.3	8.9 ± 0.4	9.2 ± 0.5	9.5 ± 0.6	10.3 ± 0.8	10 - 12
Σ^{*-}	-10.9 ± 0.2	-11.7 ± 0.2	$\textbf{-12.6}\pm0.3$	-13.1 \pm 0.3	-13.7 ± 0.3	-15.0 ± 0.4	10-12
Ξ^{*0}	8.1 ± 0.3	8.5 ± 0.3	9.0 ± 0.4	9.2 ± 0.5	9.4 ± 0.5	9.8 ± 0.7	10-12
Ξ^{*-}	-11.9 ± 0.2	-12.4 ± 0.2	-12.8 ± 0.3	-13.1 \pm 0.3	-13.4 \pm 0.3	-14.0 ± 0.3	10 - 12
Ω^{-}			-12.4 ± 0.2)			10-12

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Magnetic Polarizabilities: mesons

TABLE V. The calculated magnetic polarizabilities for the selected mesons as a function of the pion mass from the Wilson action. The pion mass is in GeV and the magnetic polarizability is in 10⁻⁴ fm³. The time window from which each polarizability is extracted is given in the last column. The errors are statistical.

κ	0.1515	0.1525	0.1535	0.1540	0.1545	0.1555	fit range
m_{π}	1.000	0.895	0.782	0.721	0.657	0.512	
π^{\pm}	-16.0 ± 0.3	-17.8 ± 0.4	$\textbf{-19.8}\pm0.5$	-21.1 \pm 0.5	-22.5 ± 0.6	-26.4 ± 0.8	16-18
π^0	8.4 ± 0.2	9.1 ± 0.3	10.1 ± 0.4	10.8 ± 0.5	11.6 ± 0.5	14.1 ± 0.8	16 - 18
K^{\pm}	-18.4 ± 0.4	-19.4 ± 0.4	-20.5 \pm 0.5	-21.1 \pm 0.5	-21.7 ± 0.6	-23.1 ± 0.6	16 - 18
K^0	3.7 ± 0.1	3.8 ± 0.2	4.0 ± 0.2	4.1 ± 0.2	4.3 ± 0.2	4.7 ± 0.3	16 - 18
ρ^{\pm}	$\textbf{-}11.8\pm0.3$	-12.5 ± 0.3	-13.1 \pm 0.4	-13.3 \pm 0.5	-13.3 \pm 0.6	-12.5 ± 1.2	14 - 16
ρ^0	5.3 ± 0.2	5.6 ± 0.2	5.9 ± 0.3	6.0 ± 0.4	6.1 ± 0.4	6.5 ± 0.7	9-11
$K^{*\pm}$	-13.0 \pm 0.3	-13.2 ± 0.4	-13.3 \pm 0.5	-13.3 \pm 0.5	-13.2 \pm 0.6	-12.7 ± 0.8	14 - 16
K^{*0}	3.0 ± 0.2	3.2 ± 0.2	3.3 ± 0.3	3.5 ± 0.3	3.6 ± 0.3	4.1 ± 0.5	14-16

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What's Next ?

Compute higher-order polarizabilities
 ←Need non-uniform, sourceless fields
 ←For example, to extract α_{E2} and γ_{E2}, choose

$$H = -\vec{\mu} \cdot \vec{B} - \frac{1}{2}\alpha E^2 - \frac{1}{2}\beta B^2$$
$$-\frac{1}{2}\gamma_{E1}\sigma \cdot \vec{E} \times \dot{\vec{E}} - \frac{1}{2}\gamma_{M1}\sigma \cdot \vec{B} \times \dot{\vec{B}}$$
$$+\gamma_{E2}\sigma_i E_{ij}B_j - \gamma_{M2}\sigma_i B_{ij}E_j$$
$$-\frac{1}{12}\alpha_{E2}E_{ij}^2 - \frac{1}{12}\beta_{M2}E_{ij}^2 + \cdots$$

$$A_{\mu} = (0, -cxt, 0, -czt), \quad \vec{E} = (cz, 0, cx)$$

For example, to extract β_{M2} and γ_{M2} , choose

$$A_{\mu} = (0, -bxy, 0, bzy), \ \vec{B} = (bz, 0, bx)$$

 $\leftarrow \alpha$ and β must be re-measured and subtracted

• The path to unquenched calculations

Use CP-PACS 2+1 flavor dynamical gauge ensembles (Iwasaki glue + clover). But still U(1) quenched

Introduce U(1) fields in the dynamical gauge generation

Conclusion

 The background field method in lattice QCD is a viable way of probing hadron internal structure

- Magnetic moments
- Electric and magnetic polarizabilities
- Neutron electric dipole moment
- Proton beta-decay
- eand more

A nice complement to experiments

Beta-decay of proton in magnetic field

 At sufficiently large B fields (10¹⁶ Tesla), proton can become heavier than neutron, allowing the 'βdecay' of the proton:

$$p \rightarrow n + e^+ + v_e$$



• As compared to the natural neutron β -decay: $n \rightarrow p + e^- + \overline{v}_e$

Such process can take place in stars where extremely strong magnetic field exists.